

# On the Extension of Adams–Bashforth–Moulton Methods for Numerical Integration of Delay Differential Equations and Application to the Moon's Orbit

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# Types of differential equations

Ordinary differential equation (ODE):

$$\dot{\mathbf{x}}(t) = \mathbf{f}(t, \mathbf{x}(t))$$

(Retarded) delay differential equation (DDE):

$$\dot{\mathbf{x}}(t) = \mathbf{f}(t, \mathbf{x}(\varphi(t)), \dots), \varphi(t) < t$$

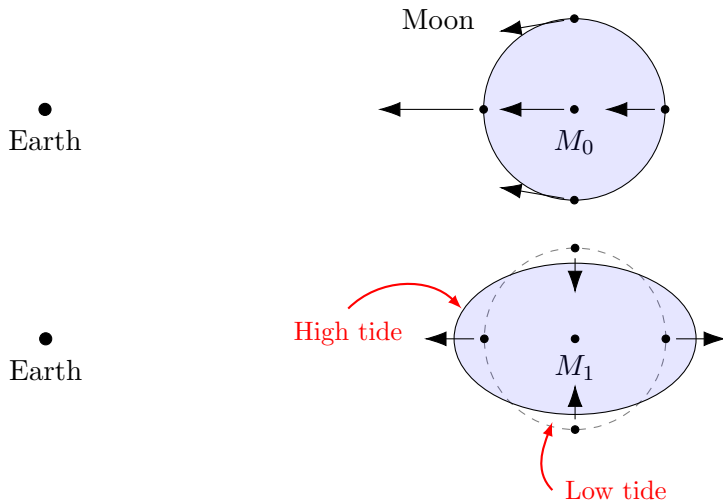
Advanced differential equation:

$$\dot{\mathbf{x}}(t) = \mathbf{f}(t, \mathbf{x}(\psi(t)), \dots), \psi(t) > t$$

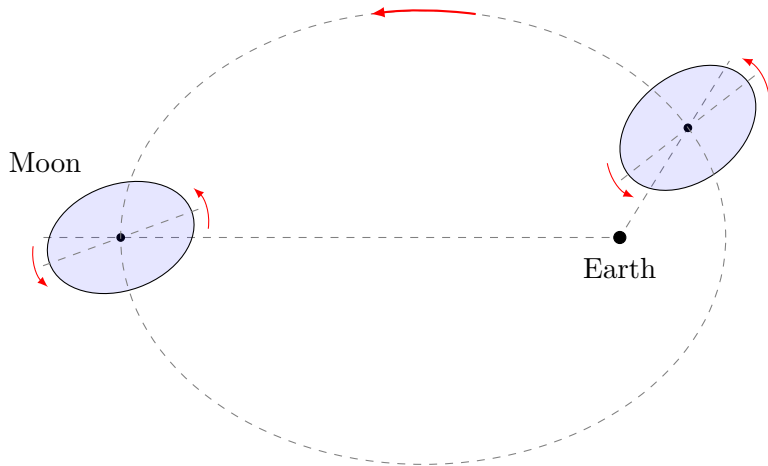
DDE of neutral type:

$$\dot{\mathbf{x}}(t) = \mathbf{f}(t, \dot{\mathbf{x}}(\xi(t)), \dots), \xi(t) \neq t$$

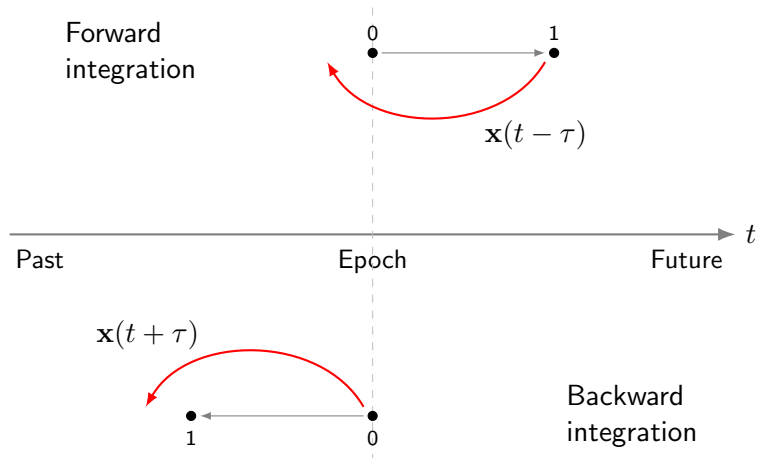
# Tidal forces (I)



## Tidal forces (II)



# From retarded to advanced equations



## The Moon equation (general form)

Forward: retarded DDE of neutral type with constant delays

$$\dot{\mathbf{x}}(t) = \mathbf{f}(t, \mathbf{x}(t), \mathbf{x}(t - \tau), \dot{\mathbf{x}}(t - \tau))$$

Backward: advanced DDE of neutral type with constant delays

$$\dot{\mathbf{x}}(t) = \mathbf{f}(t, \mathbf{x}(t), \mathbf{x}(t + \tau), \dot{\mathbf{x}}(t + \tau))$$

Initial condition at the epoch:

$$\mathbf{x}(t_0) = \mathbf{x}_0$$

## The Moon equation (actual form)

Euler's equation for a rotating reference frame:

$$\dot{\boldsymbol{\omega}} = \left( \frac{I}{m} \right)^{-1} \left[ \frac{\mathbf{N}}{m} - \frac{\dot{I}}{m} \boldsymbol{\omega} - \boldsymbol{\omega} \times \left( \frac{I}{m} \boldsymbol{\omega} \right) \right],$$

$\boldsymbol{\omega}$  — angular velocity,

$\mathbf{N}(t)$  — torque,

$I/m$  — inertia tensor

$$\frac{I}{m} = \frac{I_0}{m} - \frac{I_c}{m} - k_2 \frac{\mu_E}{\mu_M} \left( \frac{R_M}{r} \right)^5 \begin{bmatrix} x^2 - \frac{1}{3}r^2 & xy & xz \\ xy & y^2 - \frac{1}{3}r^2 & yz \\ xz & yz & z^2 - \frac{1}{3}r^2 \end{bmatrix} \\ + k_2 \frac{R_M^5}{3\mu_M} \begin{bmatrix} \omega_x^2 - \frac{1}{3}(\omega^2 - n^2) & \omega_x \omega_y & \omega_x \omega_z \\ \omega_x \omega_y & \omega_y^2 - \frac{1}{3}(\omega^2 - n^2) & \omega_y \omega_z \\ \omega_x \omega_z & \omega_y \omega_z & \omega_z^2 - \frac{1}{3}(\omega^2 - n^2) \end{bmatrix},$$

where  $\mathbf{r} = (x, y, z)^T = \mathbf{r}(t - \tau)$ ,  $\boldsymbol{\omega} = \boldsymbol{\omega}(t - \tau)$ ,  $\tau = 0.096$  d

# Runge–Kutta methods

General form for the Runge–Kutta family of methods:

$$\mathbf{x}_{n+1} = \mathbf{x}_n + h \sum_{i=1}^s b_i \mathbf{k}_i$$

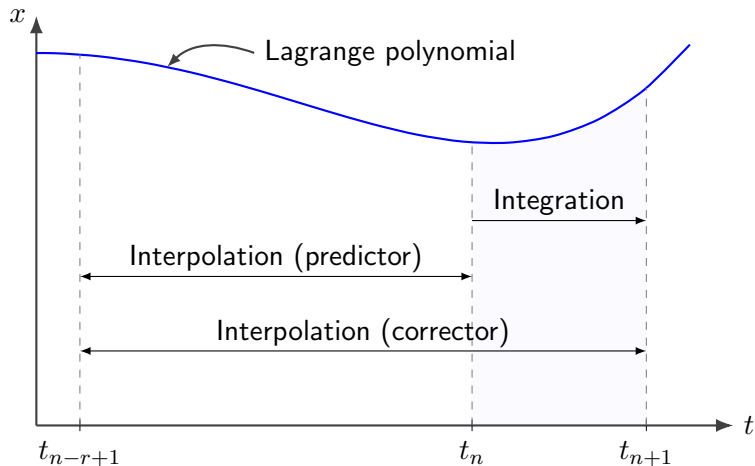
$$\mathbf{k}_s = f(t_n + c_s h, \mathbf{x}_n + h \sum_{j=1}^{s-1} a_{s,j} \mathbf{k}_j)$$

## Drawbacks

- ▶ Butcher barriers:
  - $p \geq 5$ : no RK method exists of order  $p$  with  $s = p$  stages
  - $p \geq 7$ : no RK method exists of order  $p$  with  $s = p + 1$  stages
  - $p \geq 8$ : no RK method exists of order  $p$  with  $s = p + 2$  stages
- ▶ Higher orders are problematic



# Adams–Bashforth–Moulton methods (I)



## Adams–Bashforth–Moulton methods (II)

1. Predictor — Adams–Bashforth (order 2):

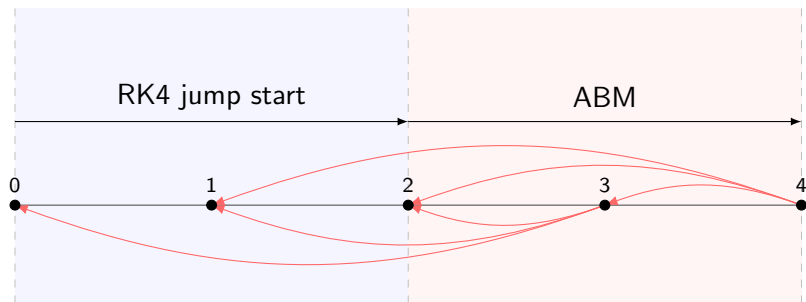
$$\mathbf{x}_{n+2} = \mathbf{x}_{n+1} + h \left( \frac{3}{2} \mathbf{f}_{n+1} - \frac{1}{2} \mathbf{f}_n \right)$$

2. Evaluation of  $\mathbf{f}_{n+2} = \mathbf{f}(t_{n+2}, \mathbf{x}_{n+2})$
3. Corrector — Adams–Moulton (order 3):

$$\mathbf{x}_{n+2} = \mathbf{x}_{n+1} + h \left( \frac{5}{12} \mathbf{f}_{n+2} + \frac{2}{3} \mathbf{f}_{n+1} - \frac{1}{12} \mathbf{f}_n \right)$$

4. (Optional). PECE, PECEC, PECECE

## Adams–Bashforth–Moulton methods (III)



First  $(r - 1)$  steps must be performed by a single-step method.

# The 'embedded RK4' method for DDEs

$$\dot{\mathbf{x}}(t) = \mathbf{f}(t, \mathbf{x}(t), \mathbf{x}(t \pm \tau), \dot{\mathbf{x}}(t \pm \tau))$$

$$\mathbf{x}(t_0) = \mathbf{x}_0$$

1. Introduce a new function

$$\mathbf{g}(t) = \mathbf{f}(t, \mathbf{x}(t), \mathbf{x}(t), \dot{\mathbf{x}}(t))$$

2. Retrieve delayed states by integrating  $\mathbf{g}(t)$  with RK4

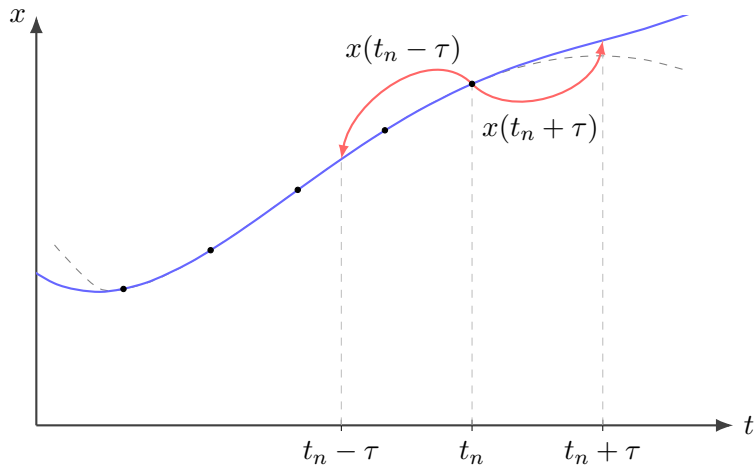
$$\mathbf{x}(t \pm \tau) = {}^{\text{RK4}}\mathcal{I}_{t \rightarrow t \pm \tau} \mathbf{g}(t)$$

$$\dot{\mathbf{x}}(t \pm \tau) = \mathbf{g}(t \pm \tau)$$

## Drawbacks

- ▶ Calculation of each delayed state requires 4 RHS calls
- ▶ No previous knowledge of  $\mathbf{x}(t)$  is being used

# The interpolation method for DDEs (I)



# The interpolation method for DDEs (II)

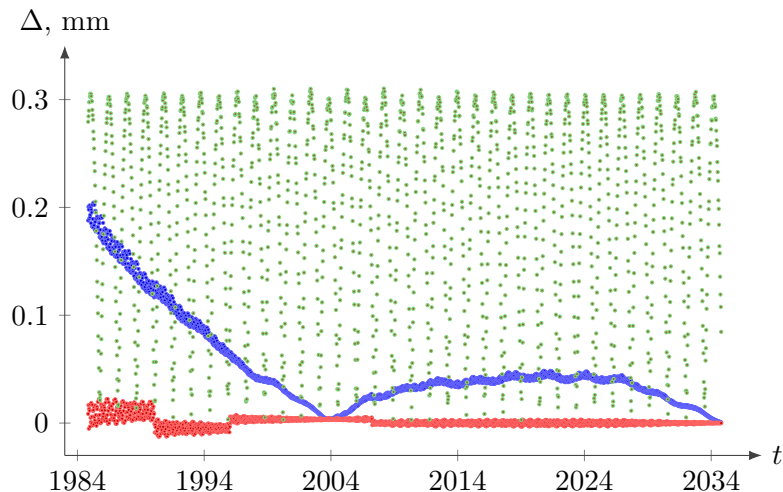
## The algorithm

1. Jump start by the 'embedded RK4' algorithm
2. P and C stages are simple ABM
3. Each E stage constructs a Lagrange interpolating polynomial (any order), which is used to find  $\mathbf{x}(t \pm \tau)$  and  $\dot{\mathbf{x}}(t \pm \tau)$

## Advantages

- ▶ Much cheaper delayed states
- ▶ No simplifying assumptions about the function

## Results. The forward-backward test



● M position ● E-M distance ● Angular displacement

Orders: ABM — 11, interpolation — 7, extrapolation — 9