



# Symbolic Sequences in the Analysis of Trajectories of Triple Black Holes

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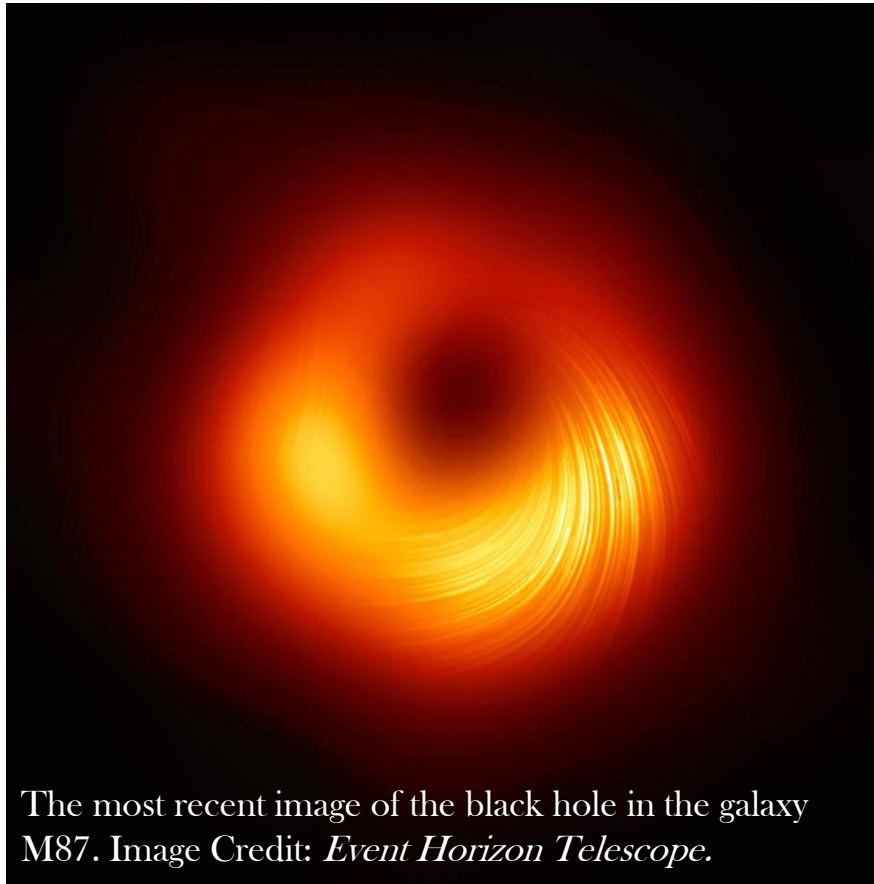
ARIEL CHITAN<sup>1</sup>, ALEKSANDR MYLLÄRI<sup>2</sup>, SHIRIN HAQUE<sup>1</sup>

POLYNOMIAL COMPUTER ALGEBRA '2021

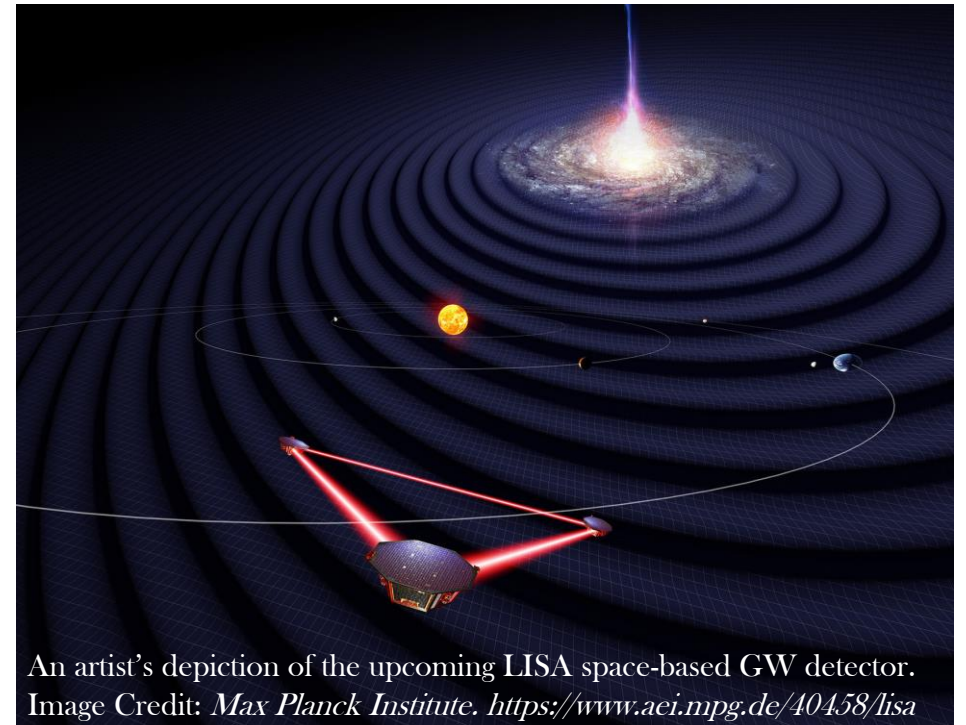


# Black Holes are thought to be abundant throughout the universe.

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We are becoming equipped with technology to detect and even image black holes.



# Triple Black Hole Systems have been observed.

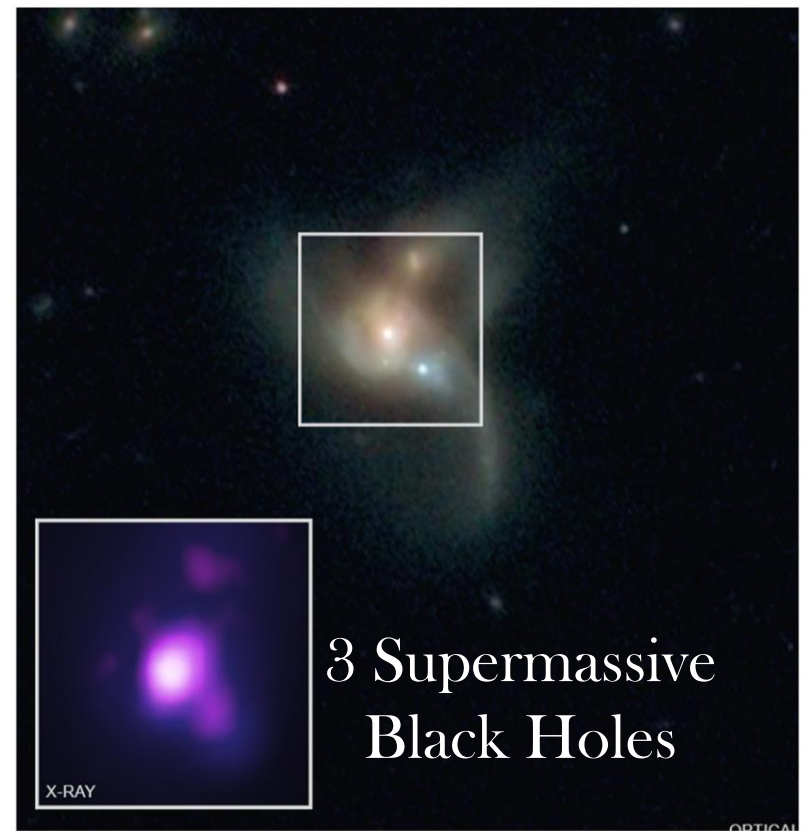


Ultraluminous Galaxy Merger NGC 6240

Spitzer Space Telescope • IRAC  
Hubble Space Telescope • ACS

NASA / JPL-Caltech / STScI/ESA / S. Bush (Harvard-Smithsonian CfA) ssc2009-06a

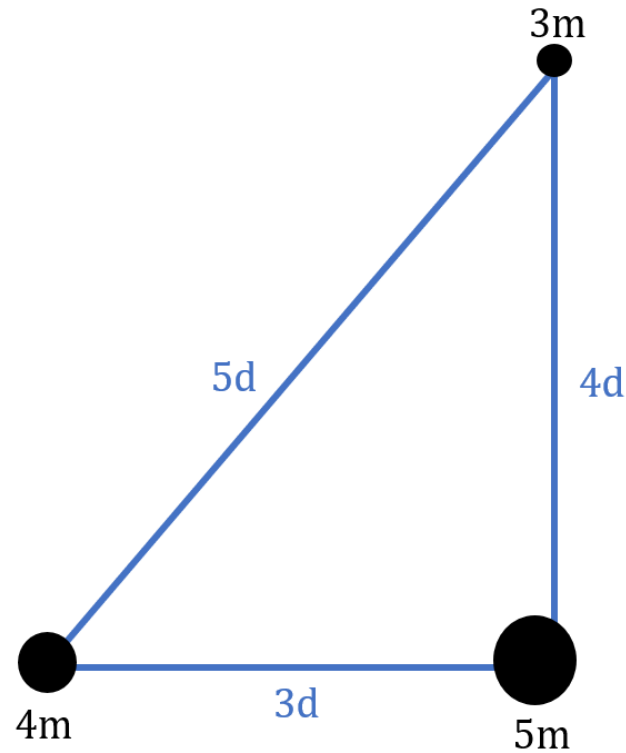
NGC  
6240 is an  
example  
of a  
merger of  
three  
galaxies.



3 Supermassive  
Black Holes

# Burrau's Problem of Three Bodies [1]

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The classical 3,4,5 system studied by Burrau (1913).

Generalised by Valtonen et al. in 1995 [2]

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**Burrau's three-body problem in the post-Newtonian approximation**

M. J. Valtonen, S. Mikkola and H. Pietilä

*Turku University Observatory, University of Turku, Tuorla, FIN-21500 Piikkiö, Finland*

Redone in 1995 with the aid of computers with  
correction terms up to the 2.5<sup>th</sup> order

We extended their study where we focus on the effect of mass [4]. The following parameters were studied:

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1. The number of mergers (we found as mass increased so did the number of mergers)
2. The lifetimes of the systems (we found that as mass increased the lifetime decreased exponentially)
3. The number of binary encounters in each system

Simulations were done using code by Mikkola et al. [3].

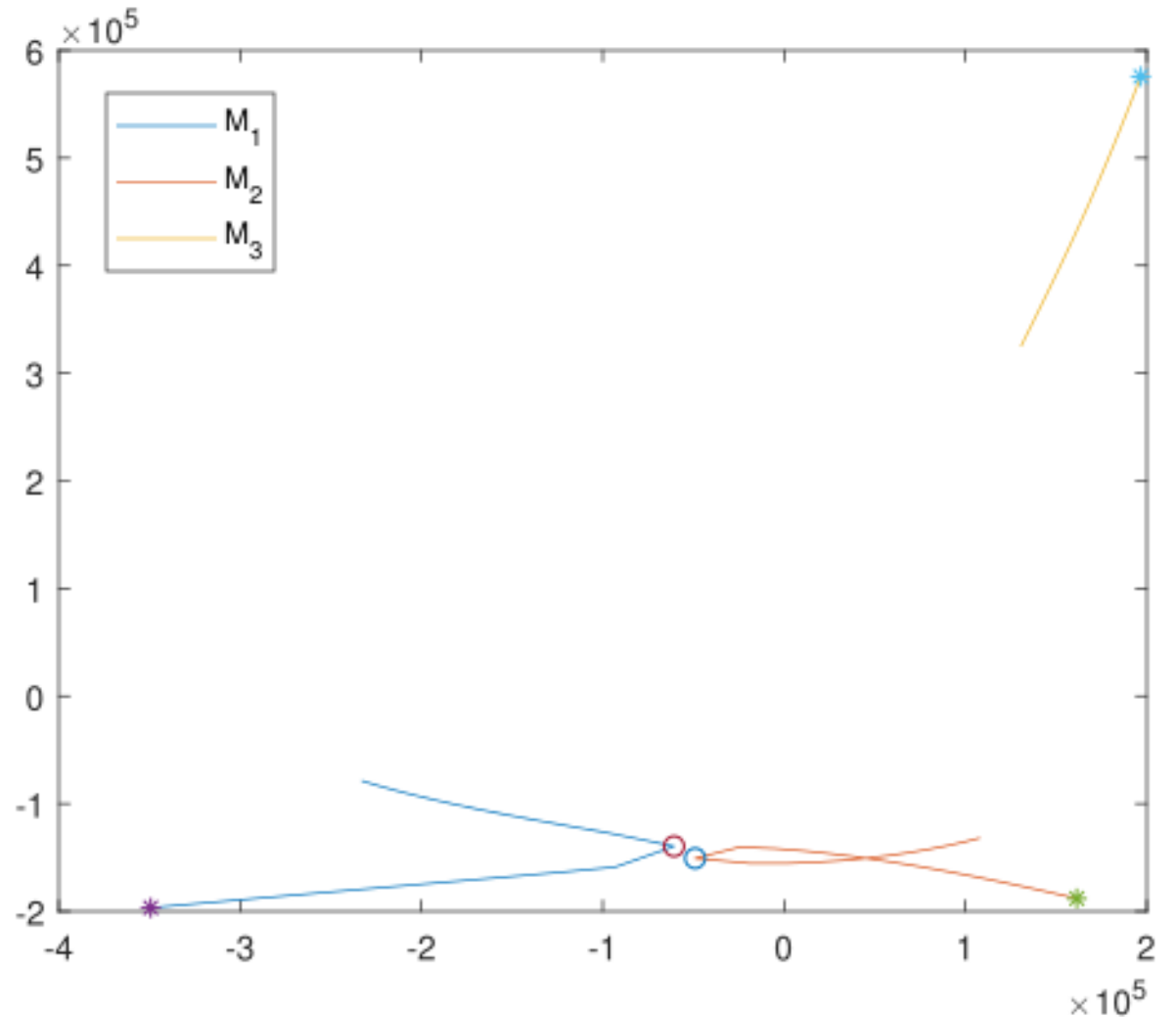
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Simulations were conducted using 2.5<sup>th</sup> order post-Newtonian equations. The code used has been updated from the one in 1995.

# Binary Encounters

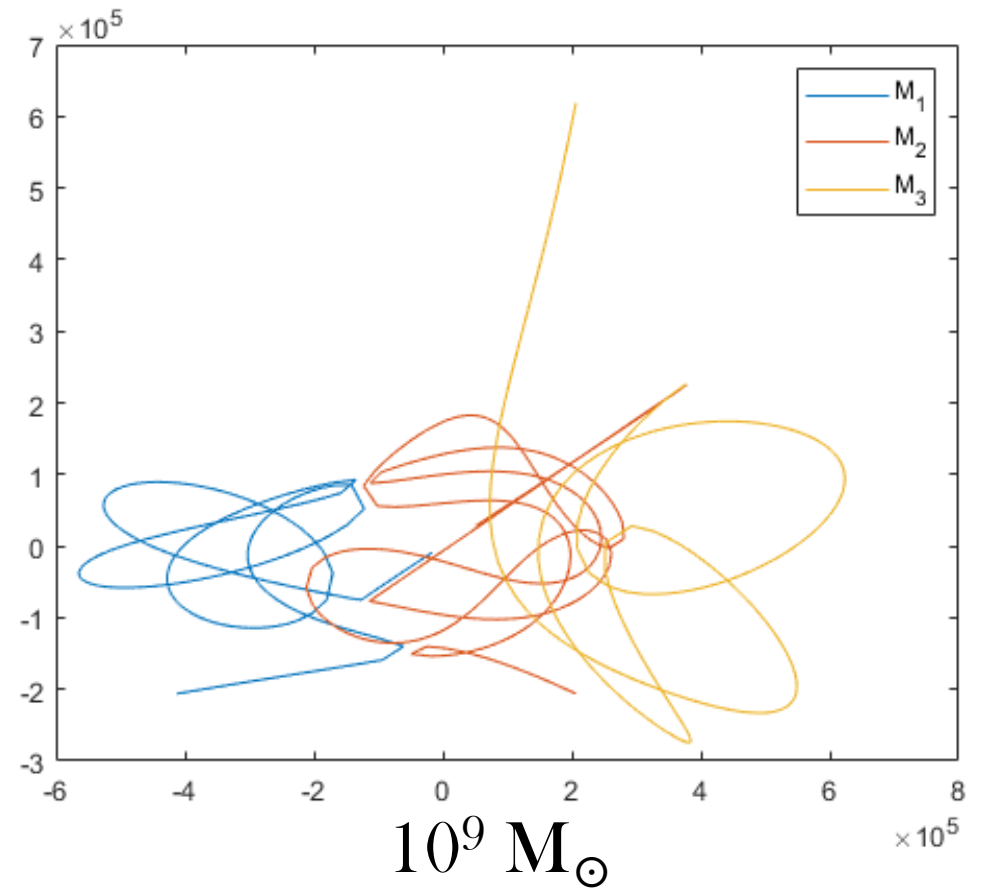
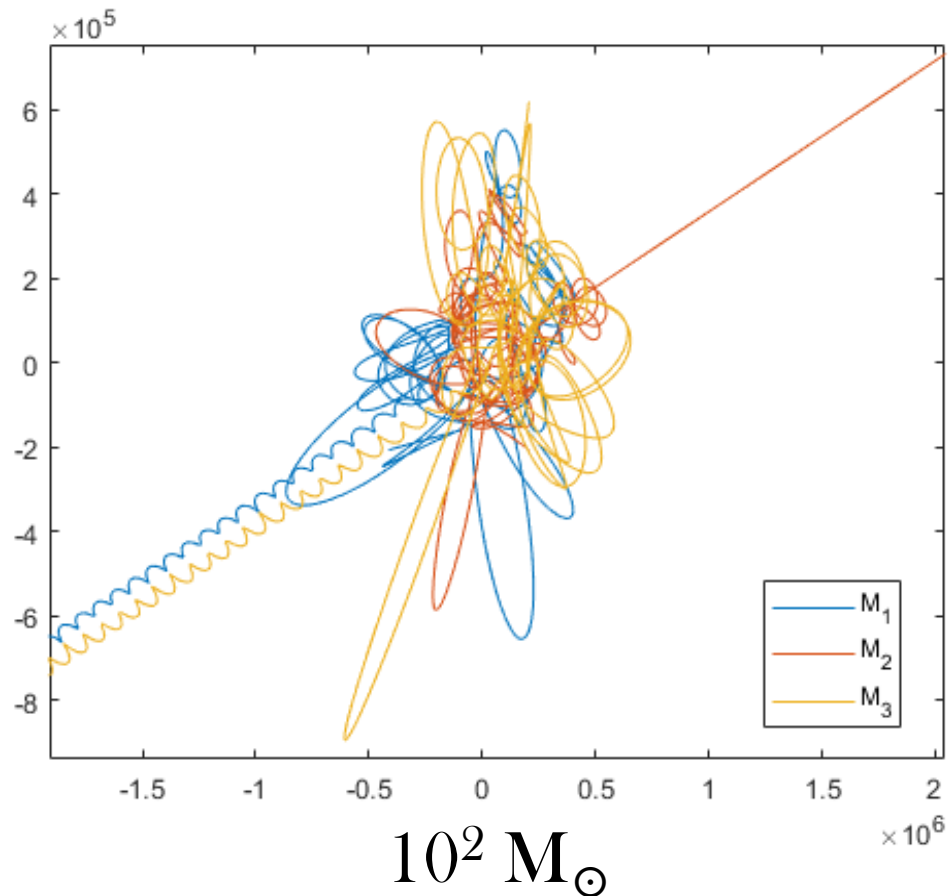
At each timestep, the distances between bodies are checked. When there is a local minimum, a binary encounter is detected.

We use this to describe how interactive systems are.

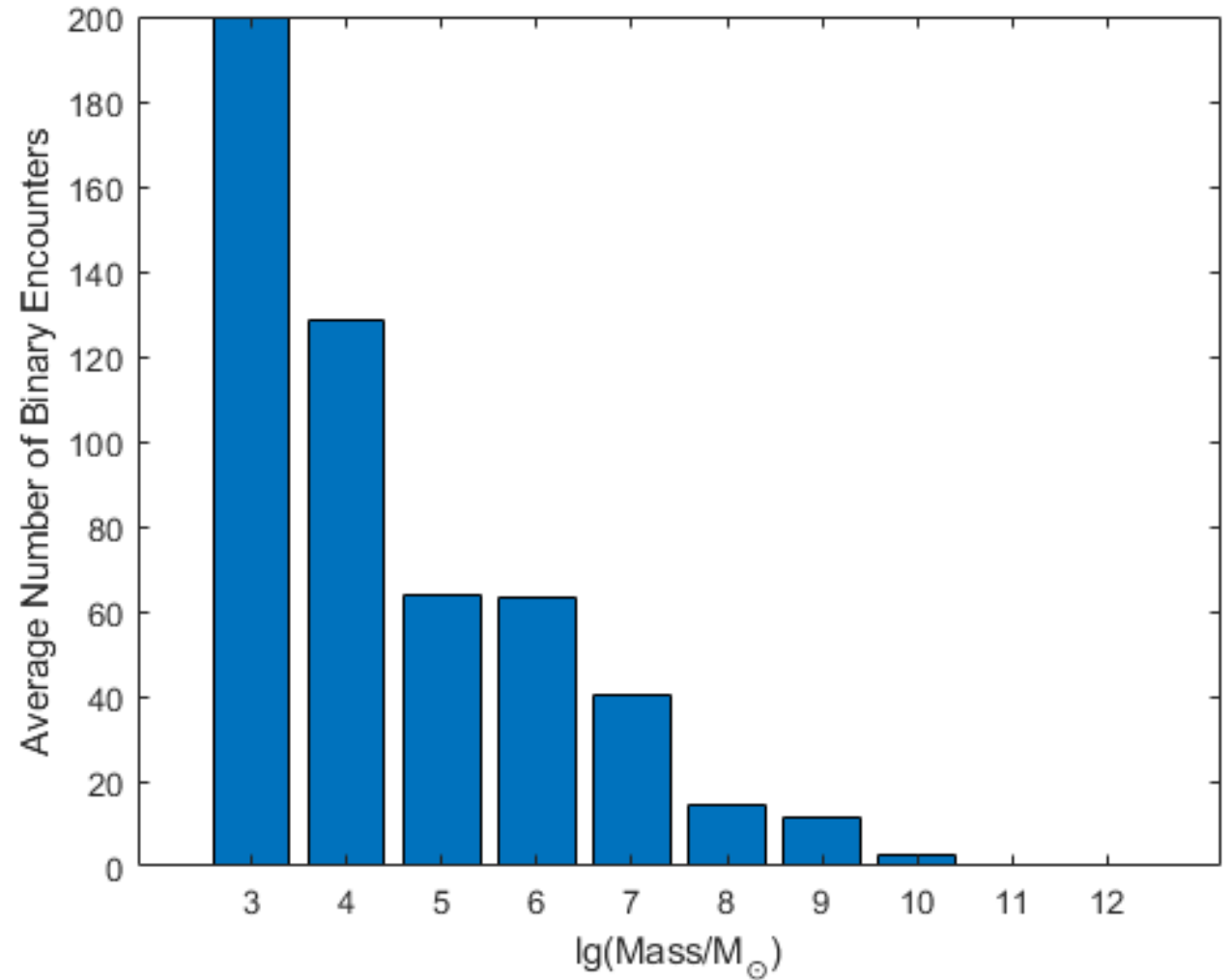




# Small mass systems tend to have more binary encounters.



Average Binary  
Encounters vs.  
Mass -  $10^3 M_{\odot}$  -  
 $10^{12} M_{\odot}$

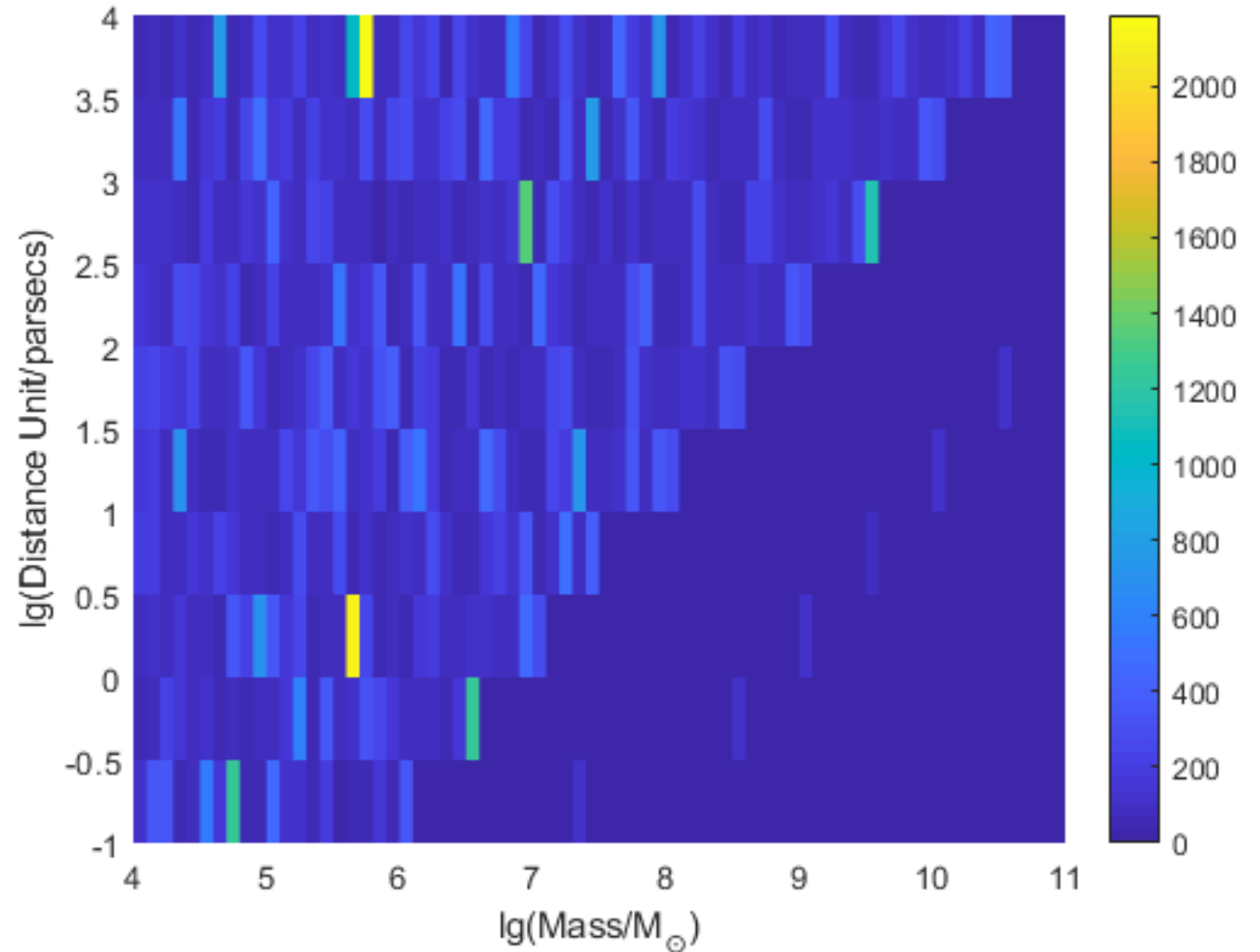


There were obvious outliers when studying the 3,4,5 system.

The number of binary encounters as both mass and distance are increased for the 3,4,5 triangle.

Here mass unit increases -  $10^4 M_{\odot}$  -  $10^{11} M_{\odot}$  and distance unit -  $10^{-1} \text{pc}$  -  $10^4 \text{pc}$ .

The Number of Binary Encounters



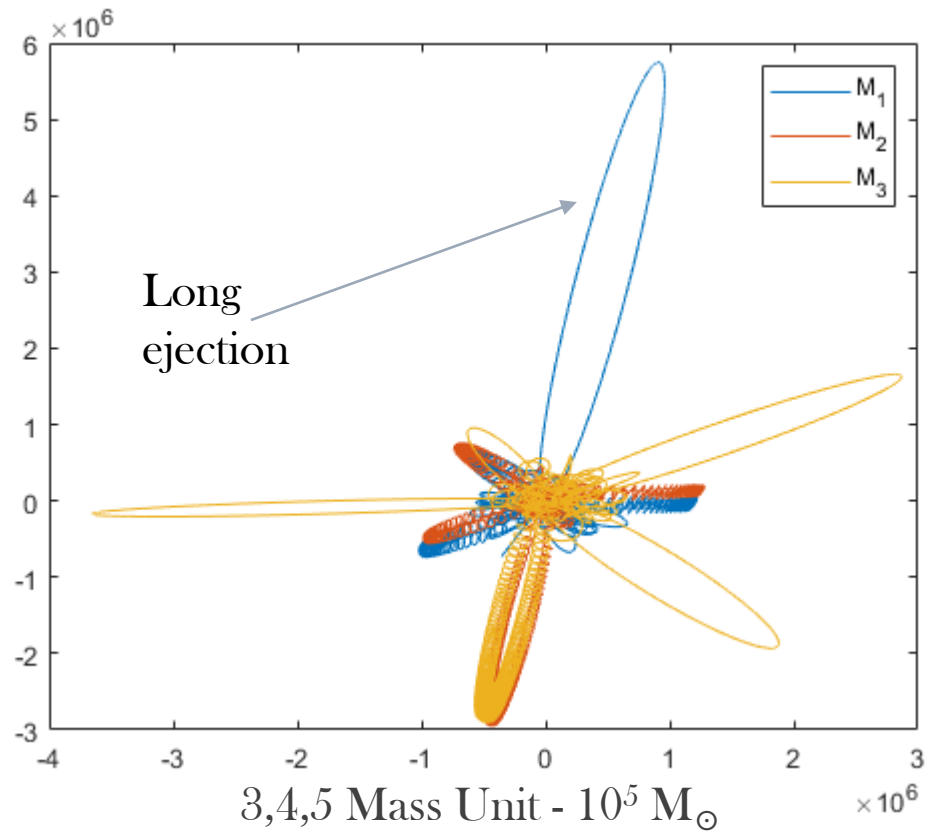
# However, there were some outliers...

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For example, the (48,55,73) triangle with a mass unit of  $10^6 M_{\odot}$  had 456 binary encounters. The average for this mass unit is around 60.

# Long ejections/third body far away or very interactive?

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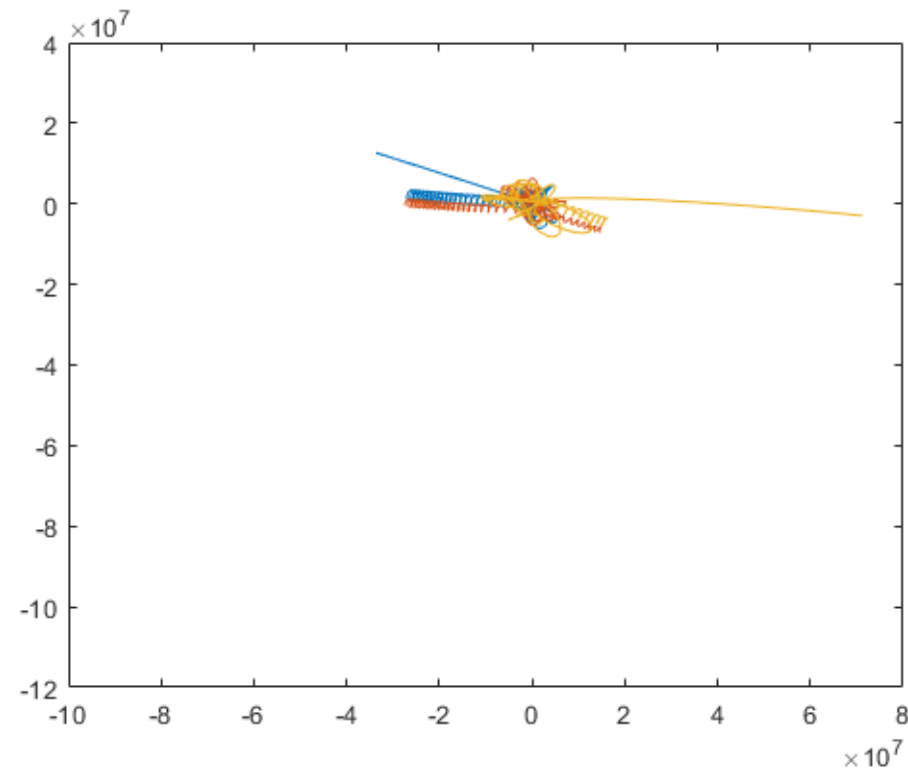


3,4,5 Mass Unit -  $10^5 M_\odot$   
- Binary Encounters = 348

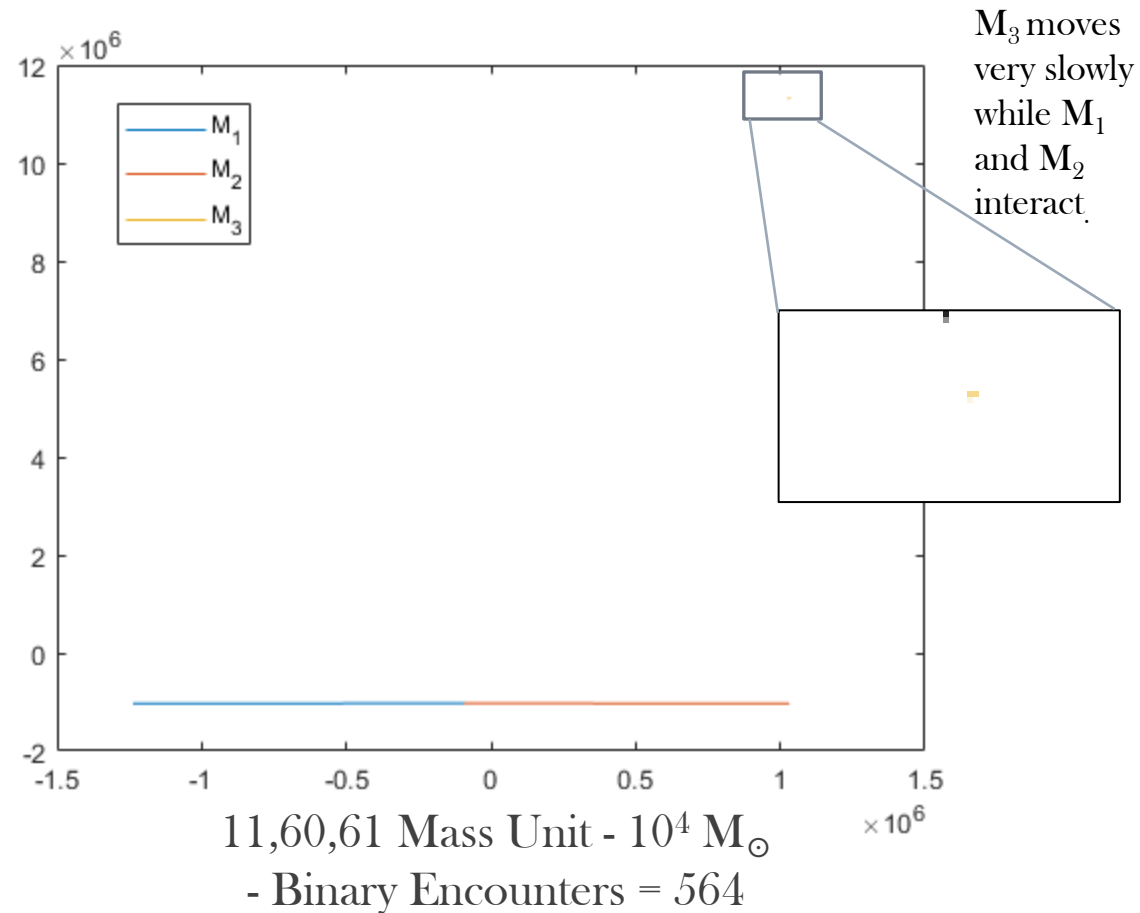


# 48,55,73 Triangle - Long Ejection

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# Or maybe the third body is far away and there is not much triple interplay



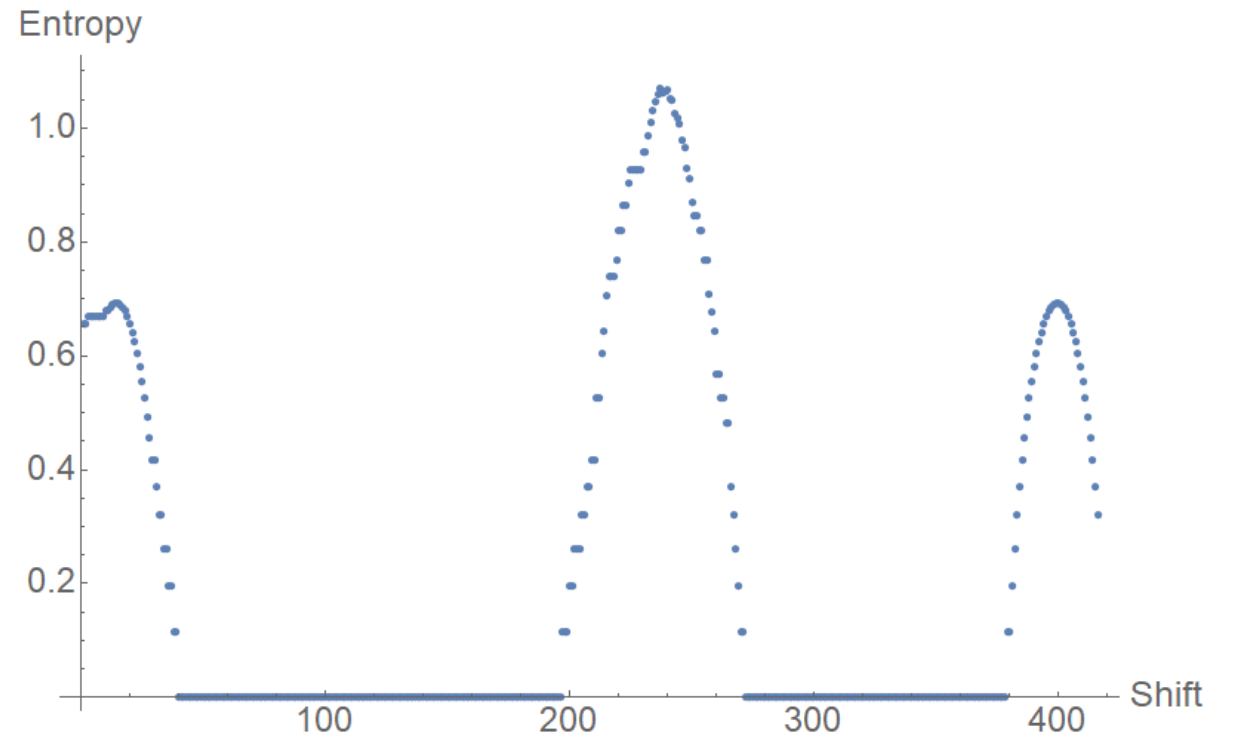
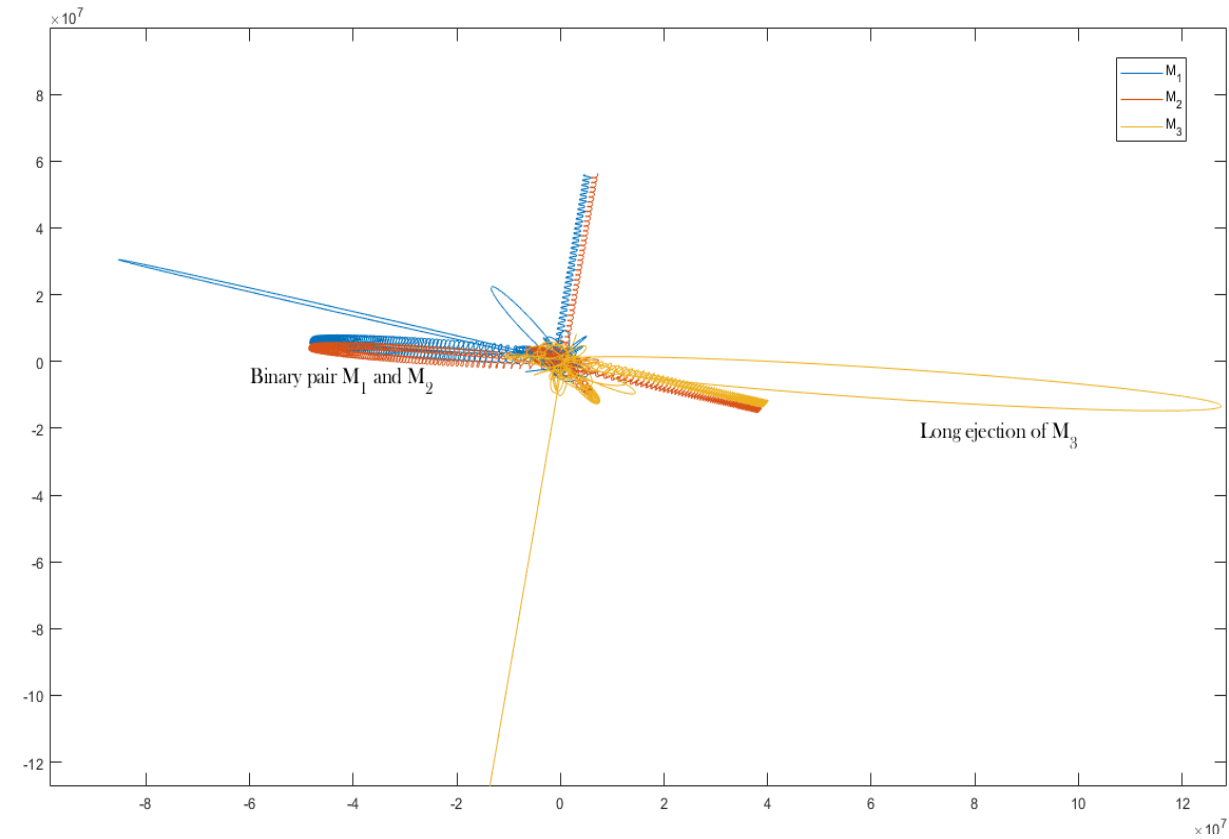
# Symbolic Sequences

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Using symbolic sequencing of the binary encounters to plot Shannon entropy helps to differentiate between these two types of systems.

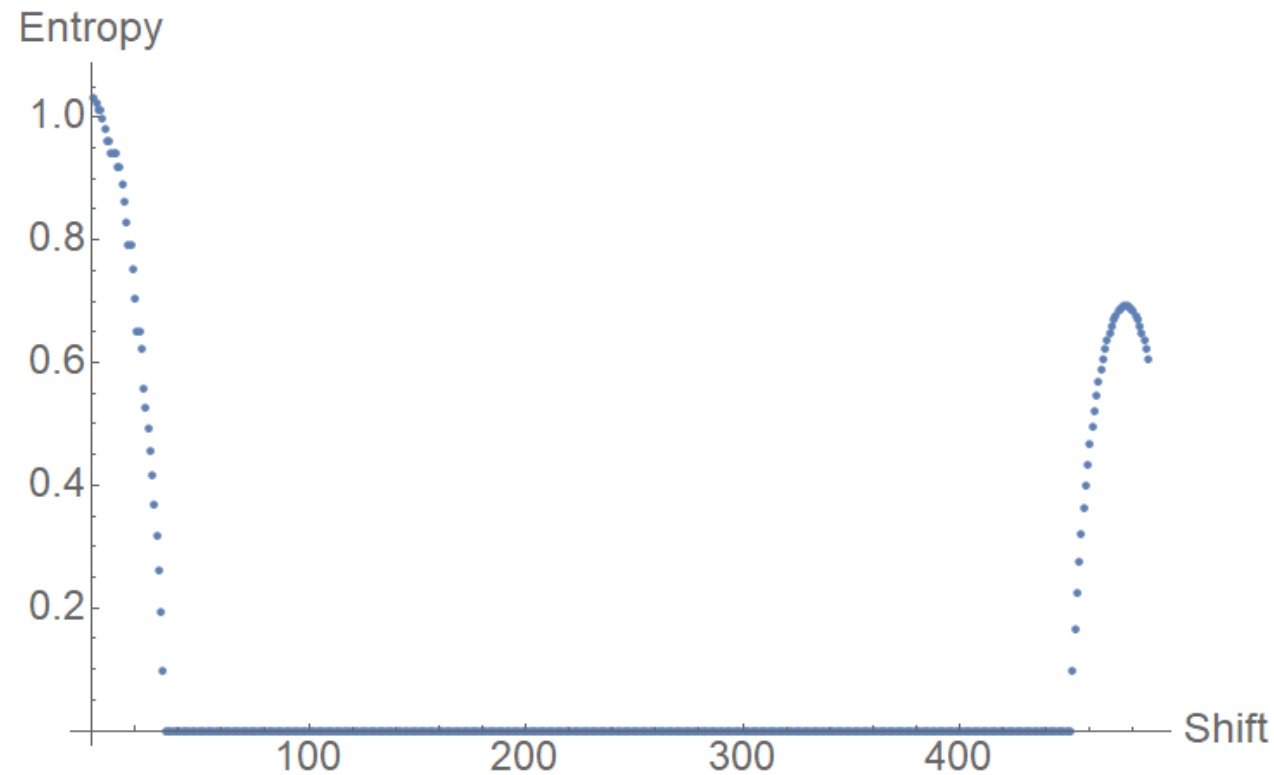
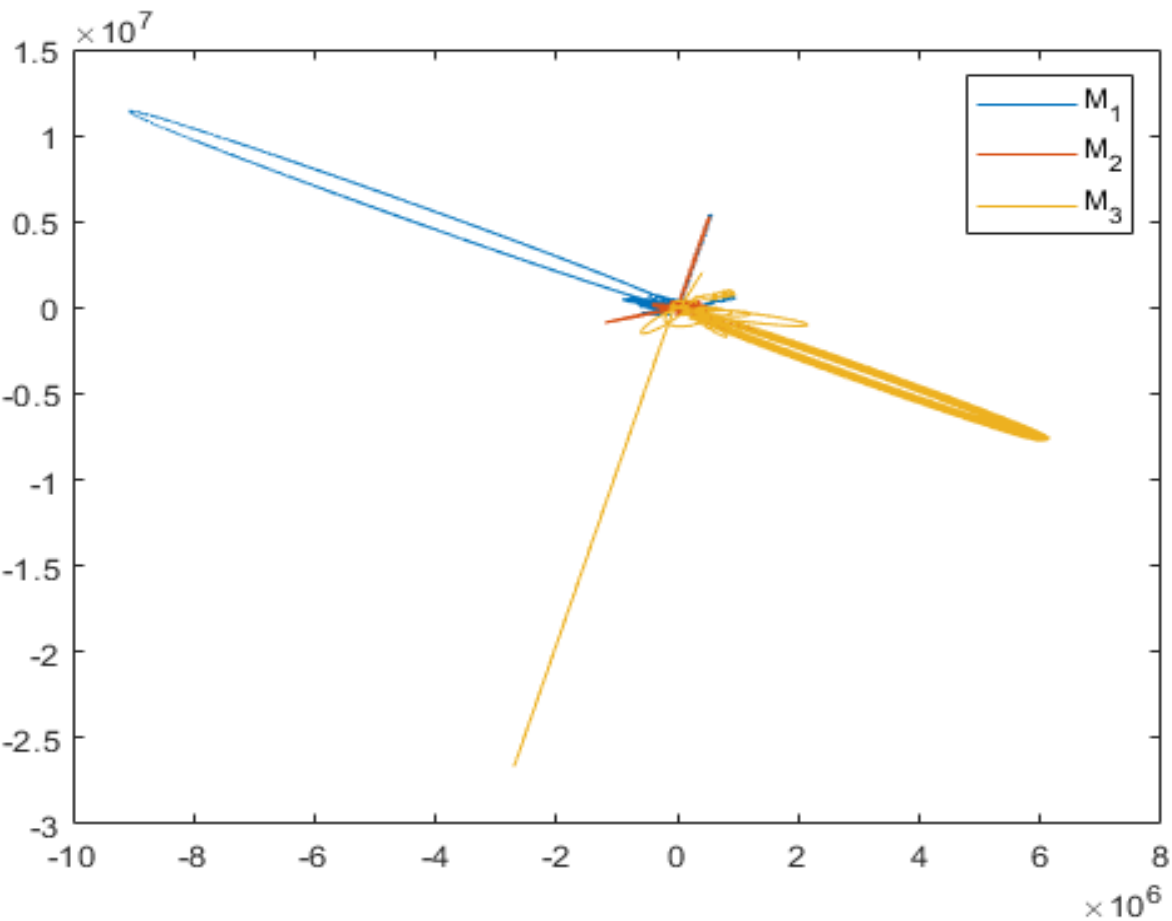
Refer to Prof. Mylläri's talk - *Free-fall three-body problem and complexity of finite symbolic sequences*. [5]

The sliding window method was discussed and is employed here as a practical example for analysing the symbolic sequences of binary encounters.



# The (48,55,73) triangle - mass unit of $10^6 M_{\odot}$

High entropy corresponds to triple interaction and interplay. Zero entropy corresponds to long ejections and binary encounters during this time between temporary binary pairs.



## The (5,12,13) triangle - mass unit of $10^4 M_{\odot}$

High entropy corresponds to triple interaction and interplay. Zero entropy corresponds to long ejections and binary encounters during this time between temporary binary pairs.



## Conclusion:

We use the symbolic sequence of binary encounters to plot Shannon entropies.

This lets us see which triples are more interactive and which have long ejections.

# References:

- [1] C. BURRAU, Numerische Berechnung eines Spezialfalles des Dreikörperproblems. *Astronomische Nachrichten* Volume 195 (Issue 6), 113–118 (1913).
- [2] M. J. VALTONEN, S. MIKKOLA, H. PIETILÄ, Burrau's three-body problem in the post-Newtonian approximation. *Monthly Notices of the Royal Astronomical Society* Volume 273 (Issue 3), 751–754 (1995).
- [3] S. MIKKOLA, K. TANIKAWA, Implementation of an efficient logarithmic-Hamiltonian three body code. *New Astronomy* Volume 20, 38–41 (2013).
- [4] A. S. CHITAN, A. MYLLÄRI, S. HAQUE, Relativistic Effects on Triple Black Holes: Burrau's Problem Revisited. *Preprint*. 2020.
- [5] A. MYLLÄRI , T. MYLLÄRI, N. VASILIEV, V. DUZHIN, Free-fall three-body problem and complexity of finite symbolic sequences, <https://pca-pdmi.ru/2021/files/13/MMVD.pdf>